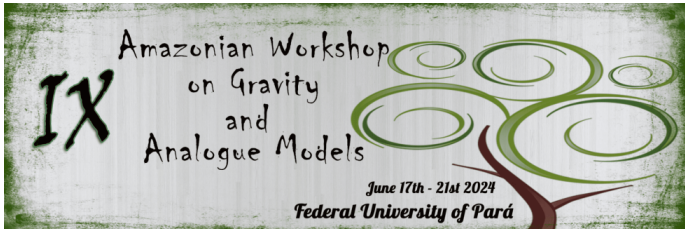


# An Analogue Model for a Particle in a Black Hole Background

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# Motivation

- Some physical systems are naturally inaccessible to us;
- However, there are systems that can mimic the properties of those inaccessible systems — we call them *analogue models*;
- Examples: The magnetic monopole as target system in spin-ice systems, position-dependent mass quantum systems, Bose-Einstein condensates etc.

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# Objectives

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- Determine the angular and radial solutions of the problem;

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# Objectives

- Solve the dynamics of a scalar particle in the background of a Kerr–de Sitter black hole;
- Determine the angular and radial solutions of the problem;
- In parallel, we investigate the dynamics of a particle constrained to move on a spheroidal surface;
- We connect both problems using an external potential, thus creating an *analogue model*.

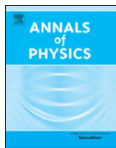


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## A quantum analog model for a scalar particle interacting with a Kerr–de Sitter black hole

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# Klein-Gordon Equation

The Kerr-de Sitter line element is  $\chi, \xi, \Xi$

$$ds^2 = -\frac{\Delta_\theta}{\Xi} Ddt^2 + \frac{r^2 + a^2 \cos^2 \theta}{r^2 + a^2} \left[ \frac{dr^2}{D} + \frac{r^2 + a^2}{\Delta_\theta} d\theta^2 \right] + \frac{r^2 + a^2}{\Xi} \sin^2 \theta d\phi^2$$

where  $D = 1 - \frac{\Lambda r^2}{3}$ ,  $\Delta_\theta = 1 + \frac{\Lambda a^2 \cos^2 \theta}{3}$ ,  $\Xi = 1 + \frac{\Lambda a^2}{3}$  and  $\Lambda$  is the cosmological constant.

# Klein-Gordon Equation

The Klein-Gordon equation can be separated using  $\psi(r, \theta, \phi, t) = F(r)G(\theta) \exp[i(m\phi + \omega t)]$ . The radial part is

$$\frac{d}{dr} \left[ (r^2 + a^2) D \frac{dF}{dr} \right] + \left( \frac{3\Xi\omega^2}{\Lambda D} - \mu^2 r^2 + \frac{m^2 a^2 \Xi}{r^2 + a^2} + \lambda \right) F = 0, \quad (1)$$

and an angular equation is

$$\frac{1}{\sin \theta} \frac{d}{d\theta} \left[ \sin \theta \Delta_\theta \frac{dG}{d\theta} \right] - \left( \frac{3\Xi\omega^2}{\Lambda \Delta_\theta} + \mu^2 a^2 \cos^2 \theta + \frac{m^2 \Xi}{\sin^2 \theta} + \lambda \right) G = 0, \quad (2)$$

# Angular Klein-Gordon

We introduce  $x = \cos^2 \theta$  into the angular equation (2),

$$G'' + \left( \frac{1}{2x} + \frac{1}{x-1} + \frac{1}{x-a_H} \right) G' + \frac{1}{4x(x-1)\Delta_x} \left[ \lambda - \frac{m^2\Xi}{x-1} + M^2 a^2 x - \frac{3\omega^2\Xi}{\Lambda\Delta_x} \right] G = 0, \quad (3)$$

with  $a_H = -\frac{3}{a^2\Lambda}$  and  $\Delta_x = 1 + \frac{\Lambda a^2 x}{3}$ .

# Angular Klein-Gordon

Applying a  $F$ -homotopic transformation

$$G(x) = x^A(x-1)^B(x-a_H)^C W(x), \quad (4)$$

with characteristic exponents equal to  $A = \{0, \frac{1}{2}\}$ ,  $B = \{\pm \frac{m}{2}\}$ , and  $C = \{\pm \frac{\omega}{2} \sqrt{\frac{3}{\Lambda}}\}$ , we get a Heun differential equation

$$\frac{d^2 y}{dx^2} + \left( \frac{\gamma}{x} + \frac{\delta}{x-1} + \frac{\varepsilon}{x-a_H} \right) \frac{dy}{dx} + \frac{\alpha\beta x - q}{x(x-1)(x-a_H)} y = 0, \quad (5)$$

## Angular Klein-Gordon

$$G_{nm}^{(1)}(\theta) = (\cos^2 \theta)^{1/2} (\cos^2 \theta - 1)^{m/2} (\cos^2 \theta - a_H)^C W_{nm}^{(1)}(\theta), \quad (6)$$

with

$$W_{nm}^{(1)}(\theta) = z^{1-\gamma} \left( \frac{z - a_H}{1 - a_H} \right)^{-\alpha - \beta + \gamma + \delta} H\ell(a_{new}, \mathbf{q}_1; \alpha_1, \beta_1, \gamma_1, \delta_1; \mathbf{z}),$$

and

$$G_{nm}^{(2)}(\theta) = (\cos^2 \theta)^{1/2} (\cos^2 \theta - 1)^{m/2} (\cos^2 \theta - a_H)^C W_{nm}^{(2)}(\theta), \quad (7)$$

with

$$W_{nm}^{(2)}(\theta) = z^{1-\delta} \left( \frac{z - a_H}{1 - a_H} \right)^{-\alpha - \beta + \gamma + \delta} H\ell(a_{new}, \mathbf{q}_2; \alpha_2, \beta_2, \gamma_2, \delta_2; \mathbf{z}),$$

where  $z = \sin^2 \theta / (1 - a_H)$ .

## Radial Klein-Gordon

The radial eigenfunctions in the correct domain are

$$O_{nm}^{(1)}(r) = \left(\frac{u}{a_H}\right)^{1-\gamma_1} \left(\frac{u-1}{a_H-1}\right)^{-\beta_1+\gamma_1-1} H\ell\left(a_{new}, q_3, \alpha_3, \beta_3, \gamma_3, \delta_3; \frac{u-a_H}{a_H(u-1)}\right), \quad (8)$$

when the characteristic exponent is zero. For the other characteristic exponent  $1 - \varepsilon_1$  the eigenfunctions reads,

$$O_{nm}^{(2)}(r) = \left(\frac{u}{a_H}\right)^{1-\gamma_1} \left(\frac{u-1}{a_H-1}\right)^{\alpha_1-\delta_1-1} (u-a_H)^{-\alpha_1-\beta_1+\gamma_1+\delta_1} \times H\ell\left(a_{new}, q_4, \alpha_4, \beta_4, \gamma_4, \delta_4; \frac{u-a_H}{a_H(u-1)}\right), \quad (9)$$

with  $u = -r^2/a^2$ .



## Radial Klein-Gordon

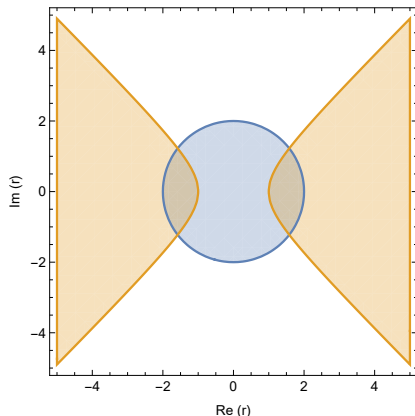


Figure: Complex  $r$ -plane. The eigenfunctions (8) and (9) converge whenever the variable  $r$  belongs to the yellow region defined by  $\frac{u-a_H}{a_H(u-1)} < \min(1, |a_{new}|)$ . Here  $u = -r^2/a^2$ , we choose  $a = 2, \Lambda = 0.5$ , so  $a_H = -3/2$  and  $a_{new} = -2/3$ .

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# Thin-Layer

- The Thin-Layer Method allows us to constrain a particle on a curved surface;
- Such curved surfaces are feasible in a lab with current technology;
- The so called *geometric potential* binds the particle on the surface. It reads

$$V_g = \frac{-\hbar^2}{2m}(M^2 - K),$$

where  $M$  and  $K$  are the mean and Gaussian curvatures;

- We can study dynamics of such particle and connect it with the dynamics of a particle in the black hole background.

# Oblate Spheroid Coordinate System

The oblate spheroidal coordinates  $(\eta, \theta, \phi)$  are defined by,

$$\begin{cases} x' = a_0 \cosh \eta \sin \theta \cos \phi \\ y' = a_0 \cosh \eta \sin \theta \sin \phi \\ z' = a_0 \sinh \eta \cos \theta \end{cases}, \quad (10)$$

where  $\eta > 0$  is a radial-like coordinate,  $\pm a_0$  are the location of the foci, is  $0 \leq \theta \leq \pi$  is the polar angle and  $0 \leq \phi \leq 2\pi$  is the azimuthal angle.

# Particle on an Oblate Spheroid

Following the thin-layer methodology [12, 13], Schrödinger equation is

$$-\frac{1}{2m^*} \left[ \frac{1}{a_0^2(s^2 + \cos^2 \theta)} \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial}{\partial \theta} \right) + \frac{1}{(s^2 + 1) \sin^2 \theta} \frac{\partial^2}{\partial \phi^2} \right] \psi + U\psi = E\psi, \quad (11)$$

The geometric potential reads,

$$V_{geom}(\theta) = -\frac{s^2 \sin^4 \theta}{4a_0^2(1 + s^2)(\cos^2 \theta + s^2)^3}. \quad (12)$$

# Particle on an Oblate Spheroid

Upon substituting  $\psi(\theta, \phi) = y(\theta)e^{im\phi}$ , if we let  $z = \cos^2 \theta$  and  $y(\sqrt{z}) = w(z)$ , we can use

$$w(z) = z^A(z-1)^B(z+s^2)^C W(z). \quad (13)$$

The differential equation is transformed into,

$$w''(z) + \left( \frac{3}{2z} + \frac{1+m}{z-1} + \frac{2+\sqrt{5}}{2(z+s^2)} \right) w'(z) + \left( \frac{A_0}{z} + \frac{B_0}{z-1} + \frac{C_0}{z+s^2} \right) w(z) = 0, \quad (14)$$

# Particle on an Oblate Spheroid

Comparing with equation (0.3) of Schäfke and Schmidt [29] — or with equation (120) of Kraniotis [30] — which is a generalized form of Heun's equation and reads,

$$y''(z) + \left( \frac{1 - \mu_0}{z} + \frac{1 - \mu_1}{z - 1} + \frac{1 - \mu_2}{z - a_H} + \alpha \right) y'(z) + \left[ \frac{\beta_0 + \beta_1 z + \beta_2 z^2}{z(z - 1)(z - a_H)} \right] y(z) = 0, \quad (15)$$

# Particle on an Oblate Spheroid

The first set of local solutions is given by,

$$y_{01}(z, \mu, \alpha, \lambda) = \eta(z, \mu_0, \mu_1, \mu_2, \alpha, \lambda; a_H), \quad (16)$$

and

$$y_{02}(z, \mu, \alpha, \lambda) = z^{\mu_0} \eta(z, -\mu_0, \mu_1, \mu_2, \alpha, \lambda; a_H), \quad (17)$$

where the Schäfke's  $\eta$ -function is defined by,

$$\eta(z, \mu, \alpha, \lambda; a_H) = \sum_{k=0}^{\infty} \frac{\tau_k(\mu, \alpha, \lambda; a_H) z^k}{\Gamma(1+k-\mu_0)\Gamma(1+k)}, \quad (18)$$

$\lambda$  and  $\mu$  stands for  $(\lambda_0, \lambda_1, \lambda_2)$ , and  $(\mu_0, \mu_1, \mu_2)$ , respectively.



# Particle on an Oblate Spheroid

The first non-normalized solution reveals

$$\psi_{nm}^{(1)}(x) = (x^2)^{1/2} (x^2 - 1)^{m/2} (x^2 + s^2)^{(2+\sqrt{5})/4} \eta \left( x^2, -\frac{1}{2}, -m, -\frac{\sqrt{5}}{2}, 0, \lambda; -s^2 \right), \quad (19)$$

and the second non-normalized solution is given by

$$\psi_{nm}^{(2)}(x) = (x^2 - 1)^{m/2} (x^2 + s^2)^{(2+\sqrt{5})/4} \eta \left( x^2, \frac{1}{2}, -m, -\frac{\sqrt{5}}{2}, 0, \lambda; -s^2 \right), \quad (20)$$

where  $x = \cos \theta$ .

# Analytic Continuation

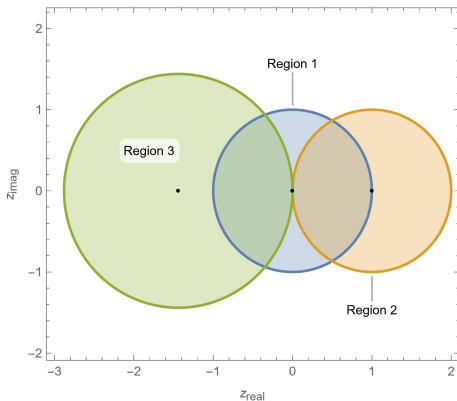


Figure: Regions of convergence of Schäfke  $\eta$  functions. The series centered at  $z = 0$  — the one we study in this work — converges inside region 1 (blue); the series centered at  $z = 1$  converges inside region 2 (yellow); and, the series centered at  $z = -s^2$  converges inside region 3 (green). Our physical domain runs from  $|z| \leq 1$ . In this figure we consider  $s = 1.2$  and the three black dots are located at  $z = 0, 1$  and  $z = -s^2$ .

# Eigenfunctions

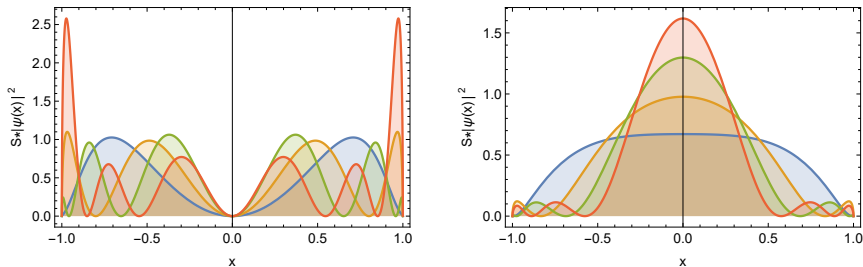


Figure: Plots of the normalized probability densities  $S|\psi(x)|^2$ , where  $x = \cos\theta$ . In the left box we plot the first solution (19) for the first four eigenenergies and  $m = 0$ . In the right box we plot the second solution (20) also for the first four eigenenergies and  $m = 0$ .

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# An Analogue Model

Carrying out a  $F$ -homotopic transformation in equation (11),

$$w(z) = z^{1/2}(z-1)^{m/2}W(z), \quad (21)$$

we need an external potential with the form

$$V_{\text{ext}}(z) = \frac{8}{a_0^2} \left[ \frac{U_0(z-1)}{(z+s^2)^3} + \frac{U_1z}{(z-1)(z+s^2)^2} + \frac{U_2z}{(z+s^2)^3} + \frac{U_3z^2}{(z-1)^2(z+s^2)} + \frac{U_4}{(z+s^2)} \right]. \quad (22)$$

# Heun's Differential Equation

The resulting equation will be of the form

$$\frac{d^2 W}{dz^2} + \left( \frac{\gamma}{z} + \frac{\delta}{z-1} + \frac{\varepsilon}{z-a_H} \right) \frac{dW}{dz} + \frac{\alpha\beta z - q}{z(z-1)(z-a_H)} W(z) = 0, \quad (23)$$

which is the general Heun equation.

# Conclusions










- Eigenenergies and eigenfunctions of the black hole problem using Heun's functions;
- Eigenenergies and eigenfunctions of the spheroidal surface problem using Schäfke's  $\eta$ -functions;
- Applying an external field, we created an analogue model for the dynamics of a particle in the background of a Kerr-de Sitter black hole;











Thank you!  
Obrigado!



































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